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REDISTRIBUTION OF VELOCITY: COLLISION TRANSFORMATIONS

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Suppose we have a large number of particles of equal mass with an initial distribution of velocity. We assume that these particles undergo triple collisions at random and that the total velocity of each triple is redistributed according to some given redistribution law. We show that for each given redistribution law there is an attractive invariant velocity distribution. There is a distribution of velocity such that for any nontrivial initial distribution with finite moments of all orders, the iterates converge weakly to this stable distribution.

This work is a natural generalizations of Ulam's redistribution of energy problem [2]. In fact, Ulam had speculated that such theorem may be true. In developing our approach to the problem, we benefitted from computer studies which strongly indicated that the result hold in some cases. We thank Tony Warnock of Cray Research for conducting these studies. The formal setting is developed in sections 1 and 2 and the Main Theorem (Theorem 2.1) is stated. Moment recursion formulas and their convergence are developed in sections 3, 4, and 5. In section 6, the proof of Theorem 2.1 is completed and a partial converse (Theorem 6.1) is proven. Finally, in section 7, the uniform redistribution law is shown to yield a normal velocity distribution (Theorem 7.3).

1. THE SETTING

Consider three particles of equal mass which form a complex, and the velocities of the particles are redistributed with the constraints that the total energy and momentum are conserved. Thus, if \mathbf{v}_1 , \mathbf{v}_2 , and \mathbf{v}_3 are the initial velocity vectors in \mathbf{R}^3 of the particles, then the new velocities \mathbf{v}'_1 , \mathbf{v}'_2 , and \mathbf{v}'_3 satisfy:

$$(1.1) \quad \mathbf{S}_1 := \mathbf{v}_1 + \mathbf{v}_2 + \mathbf{v}_3 = \mathbf{v}'_1 + \mathbf{v}'_2 + \mathbf{v}'_3 := \mathbf{S}'_1,$$

and

$$(1.2) \quad S_2 := \|\mathbf{v}_1\|^2 + \|\mathbf{v}_2\|^2 + \|\mathbf{v}_3\|^2 = \|\mathbf{v}'_1\|^2 + \|\mathbf{v}'_2\|^2 + \|\mathbf{v}'_3\|^2 := S'_2.$$

We will consider redistribution of energy and velocity in the center of mass frame of reference. Let λ_i be the fraction of the total kinetic energy that the i^{th} particle has after collision and let \mathbf{w}_i be the direction vector of the velocity of the i^{th} particle after collision. For convenience, we assume all particles have mass 1. Thus,

$$(1.3) \quad 0 \leq \lambda_i = K_i/K$$

where K is the kinetic energy measured in the center of mass frame of reference, $K = S_2 - \|\mathbf{S}_1\|^2/3$. So,

$$(1.4) \quad \lambda_1 + \lambda_2 + \lambda_3 = 1.$$

Since the total velocity in the center of mass system is zero, we have

$$(1.5) \quad \sqrt{\lambda_1} \mathbf{w}_1 + \sqrt{\lambda_2} \mathbf{w}_2 + \sqrt{\lambda_3} \mathbf{w}_3 = 0.$$

We derive another form of the constraints on the λ 's which are more suitable for our purposes. From (1.5), we get

$$(1.6) \quad \lambda_1 + 2\sqrt{\lambda_1\lambda_2} \langle \mathbf{w}_1, \mathbf{w}_2 \rangle + \lambda_2 = \lambda_3.$$

Using (1.4), we have

$$(1.7) \quad |1 - 2(\lambda_1 + \lambda_2)| \leq 2\sqrt{\lambda_1\lambda_2}.$$

After some algebra,

$$(1.8) \quad 4\lambda_1^2 - 4\lambda_1 + 4\lambda_2^2 - 4\lambda_2 + 1 + 4\lambda_1\lambda_2 \leq 0.$$

From this inequality, we derive

$$(1.9) \quad (1/3)(18\lambda_1^2 - 18\lambda_1 + 18\lambda_2^2 - 18\lambda_2 + 6 + 18\lambda_1\lambda_2) \leq 1/2.$$

Now, (1.9) can be expressed as

$$(1.10) \quad (1/3)(9\lambda_1^2 - 6\lambda_1 + 9\lambda_2^2 - 6\lambda_2 + 9(1 - \lambda_1 - \lambda_2)^2 - 6(1 - \lambda_1 - \lambda_2) + 3) \leq 1/2.$$

Using the identity (1.4), we get the inequality

$$(1.11) \quad (\lambda_1 - 1/3)^2 + (\lambda_2 - 1/3)^2 + (\lambda_3 - 1/3)^2 \leq 1/6.$$

Thus, conditions (1.4) and (1.11) imply that the point $[\lambda_1, \lambda_2, \lambda_3]$ must lie on the circular disk, D , of radius $\frac{1}{\sqrt{6}}$, center $[\frac{1}{3}, \frac{1}{3}, \frac{1}{3}]$ which lies in the plane given by equation (1.4). The mutually orthogonal vectors $[\frac{1}{3}, -\frac{1}{6}, -\frac{1}{6}]$ and $[0, \frac{1}{2\sqrt{3}}, -\frac{1}{2\sqrt{3}}]$ both have length $\frac{1}{\sqrt{6}}$, and are both orthogonal to the normal, $[1, 1, 1]$, of equation (1.4). Thus for each $\mathbf{x} \in D \setminus \{[\frac{1}{3}, \frac{1}{3}, \frac{1}{3}]\}$, there exists a unique $r \in (0, 1]$ and a unique $\theta \in [0, 2\pi)$, so that

$$(1.12) \quad \mathbf{x} = \left[\frac{1}{3}, \frac{1}{3}, \frac{1}{3}\right] + r \cos \theta \left[\frac{1}{3}, -\frac{1}{6}, -\frac{1}{6}\right] + r \sin \theta \left[0, \frac{1}{2\sqrt{3}}, -\frac{1}{2\sqrt{3}}\right].$$

Conversely, if $[\lambda_1, \lambda_2, \lambda_3]$ lies on the disk D , then tracing backwards from (1.11) (using (1.4) when necessary) it is easy to see that $[\lambda_1, \lambda_2, \lambda_3]$ also satisfies (1.7). Letting \mathbf{w}_1 and \mathbf{w}_2 denote unit vectors in \mathbf{R}^3 so that $2\sqrt{\lambda_1\lambda_2} \langle \mathbf{w}_1, \mathbf{w}_2 \rangle = 1 - 2(\lambda_1 + \lambda_2)$, the vector $[\lambda_1, \lambda_2, \lambda_3]$ must satisfy $\|\sqrt{\lambda_1}\mathbf{w}_1 + \sqrt{\lambda_2}\mathbf{w}_2\|^2 = \lambda_3$. Letting $-\mathbf{w}_3$ denote the direction vector for $\sqrt{\lambda_1}\mathbf{w}_1 + \sqrt{\lambda_2}\mathbf{w}_2$, we have $\sqrt{\lambda_1}\mathbf{w}_1 + \sqrt{\lambda_2}\mathbf{w}_2 = -\sqrt{\lambda_3}\mathbf{w}_3$. Thus all points in this disk, D , are realizable as values of λ_1, λ_2 and λ_3 which satisfy (1.4) and (1.5) for some set of unit vectors.

A redistribution of energy law is a probability measure $\tilde{\mu}$ supported on the disk D which is symmetric, or invariant under permutations of the coordinates. This last condition signifies that the particles are indistinguishable. The direction vectors for the velocity of the particles are chosen, independent of $\tilde{\mu}$, as follows. First, \mathbf{w}_1 is chosen from the unit sphere according to the uniform distribution. Next, a unit vector \mathbf{z} which is perpendicular to \mathbf{w}_1 is chosen according to the uniform distribution on the great circle which is the intersection of the unit sphere and the plane normal to \mathbf{w}_1 . Thus, \mathbf{w}_1 and \mathbf{z} determine a plane which contains \mathbf{w}_2 and \mathbf{w}_3 . Finally, \mathbf{w}_2 and \mathbf{w}_3 are determined up to the reflection $\mathbf{y} \rightarrow -\mathbf{y}$ in this plane from equations (1.4), (1.5), and the equality $\|\mathbf{w}_2\| = \|\mathbf{w}_3\| = 1$.

We will study this process under iteration. We note that it is immaterial whether we chose \mathbf{w}_1 at random first. Thus, we have not violated the indistinguishability of the particles. Also, we note that there are 9 velocity component variables which

have 5 degrees of freedom in view of the conservation laws. In our scheme, we have 2 degrees of freedom in choosing the λ 's, 2 degrees in choosing \mathbf{w}_1 and one degree in choosing \mathbf{z} , a total of five.

2. THE REDISTRIBUTION OPERATOR

We now formalize the redistribution operator, $T_{\tilde{\mu}}$. Let ν be a probability measure on \mathbf{R}^3 and assume we have a vast number of particles with velocity distribution ν . We imagine that these particles are partitioned into triples at random. For each triple the velocity is redistributed as described in section one, which yields a new velocity distribution, $T_{\tilde{\mu}}(\nu)$. So, if $\mathbf{X}_1, \mathbf{X}_2,$ and \mathbf{X}_3 are independent random velocity vectors each distributed as ν , then $T_{\tilde{\mu}}(\nu)$ will be the distribution of \mathbf{X}'_1 , since $\mathbf{X}'_1, \mathbf{X}'_2, \mathbf{X}'_3$ all have the same distribution.

From this point on, we will suppress the subscript $\tilde{\mu}$ in $T_{\tilde{\mu}}$. Thus, $T(\nu)$ is the distribution of

$$\frac{\mathbf{S}_1}{3} + \sqrt{K\lambda}\mathbf{u},$$

where (1) \mathbf{u} has the uniform distribution, π_2 , on the unit sphere, (2) λ is distributed as $\mu = \tilde{\mu} \circ p^{-1}$ on $[0, 2/3]$, where $\tilde{\mu}$ is the redistribution law on the disk D and p is the projection map of \mathbf{R}^3 onto the first coordinate, and, (3) K is the kinetic energy in the center of mass frame of reference; *i.e.*, $K = S_2 - \|\mathbf{S}_1\|^2/3$. Therefore, $T(\nu)$ is determined by the functional equation:

$$(2.1) \quad T\nu(f) = \int_{\mathbf{R}^3} f(\mathbf{x}) dT\nu(\mathbf{x}) =$$

$$\int \dots \int f \left(\frac{\sum_{i=1}^3 \mathbf{x}_i}{3} + \sqrt{\lambda} \sqrt{\sum_{i=1}^3 \|\mathbf{x}_i\|^2 - \frac{\left\| \sum_{i=1}^3 \mathbf{x}_i \right\|^2}{3}} \mathbf{u} \right) d\nu(\mathbf{x}_1) d\nu(\mathbf{x}_2) d\nu(\mathbf{x}_3) d\pi_2(\mathbf{u}) d\mu(\lambda),$$

where the integral is over $[0, \frac{2}{3}] \times S^2 \times \mathbf{R}^3 \times \mathbf{R}^3 \times \mathbf{R}^3$.

We will abbreviate integrals like this last one unless the domain of integration or integrators are not clear. Thus, (2.1) could be written as

$$(T\nu)(f) = \int \dots \int f \left(\frac{\mathbf{S}_1}{3} + \sqrt{\lambda} \sqrt{S_2 - \frac{\|\mathbf{S}_1\|^2}{3}} \mathbf{u} \right) d\nu d\nu d\nu d\pi_2 d\mu.$$

We note a few basic properties of the non-linear operator $T : Prob(\mathbf{R}^3) \rightarrow Prob(\mathbf{R}^3)$. First, T is weakly continuous and commutes with the translation operators:

$$(2.2) \quad T(\nu(\cdot + \mathbf{x}_0)) = T\nu(\cdot + \mathbf{x}_0),$$

for any $\mathbf{x}_0 \in \mathbf{R}^3$. This follows from using (2.1) and the fact that the function $\sqrt{S_2 - \|\mathbf{S}_1\|^2/3}$ is invariant under translation in \mathbf{R}^3 . The operator T preserves energy:

$$(2.3) \quad \nu(\|\mathbf{x}\|^2) = T\nu(\|\mathbf{x}\|^2)$$

$$K = \sum_{i=1}^3 \|\mathbf{x}_i - \bar{\mathbf{x}}\|^2$$

and momentum:

$$(2.4) \quad \nu(\mathbf{x}) = T\nu(\mathbf{x}).$$

Equations (2.3) and (2.4) can be verified by using (2.1) and the facts that

$$(2.5) \quad \int_D x_1 d\tilde{\mu}(\mathbf{x}) = \int_{[0,2/3]} \lambda d\mu(\lambda) = 1/3,$$

and

$$(2.6) \quad \int_{S^2} u_i(\mathbf{u}) d\pi_2(\mathbf{u}) = 0,$$

where u_i is the i^{th} coordinate of \mathbf{u} .

In order to see that (2.5) holds, we will use Choquet's representation theorem [3]. We will also use this representation later on. Our measure $\tilde{\mu}$ is simply a probability measure on the disk D which is invariant under the symmetries of the circumscribing triangle with vertices $\mathbf{e}_1 = [1, 0, 0]$, $\mathbf{e}_2 = [0, 1, 0]$, and $\mathbf{e}_3 = [0, 0, 1]$. Therefore, $\tilde{\mu}$ can be expressed uniquely as an integral over the set of all extreme points of \mathbf{C} , the compact convex set of all probability measures which are invariant under these symmetries. Thus, there is a unique probability measure P on $\text{ext}(\mathbf{C})$ such that

$$(2.7) \quad \tilde{\mu}(f) = \int_{\text{ext}(\mathbf{C})} \int_D f(\mathbf{x}) d\tau(\mathbf{x}) dP(\tau).$$

Now, an extreme point in this case is simply a probability measure on D which is ergodic under the action of the symmetry group. It is easy to see that τ is ergodic if and only if there is a point \mathbf{z} of D such that

$$(2.8) \quad \tau = \frac{1}{6} \sum_{i=0}^5 \delta_{\xi_i(\mathbf{z})},$$

where $\xi_0, \xi_1, \dots, \xi_5$ are the symmetries and δ_x is point mass at x .

Of course, for each such τ ,

$$(2.9) \quad \int_D x_1 d\tau(\mathbf{x}) = 1/3.$$

Since $\int_{[0,2/3]} \int_D g d\mu = \int_D g \circ pd\tilde{\mu}$, (2.5) follows from (2.9).

Also, since the function $\sqrt{S_2 - \|\mathbf{S}_1\|^2/3}$ is positive homogeneous, T commutes with positive scaling:

$$(2.10) \quad T(\nu(c \cdot)) = T\nu(c \cdot).$$

Facts concerning weak convergence of probability measures in metric spaces such as \mathbf{R}^3 can be found in [1]. Our main theorem concerns the properties of the fixed point and its domain of attraction.

Theorem 2.1 (Main). *For each symmetric probability measure $\tilde{\mu}$, on D , there exists a unique radially symmetric probability measure $\hat{\mu}$ on \mathbf{R}^3 with total energy one and moments of all orders so that $\hat{\mu} = T_{\tilde{\mu}}(\hat{\mu})$. Further, if ν is any probability measure on \mathbf{R}^3 with finite moments of all orders and ν is not point mass, then*

$$(2.11) \quad \{T^n(\nu)\}_{n=0}^{\infty} \text{ converges weakly to } \hat{\mu} \left[\frac{\cdot - \nu(\mathbf{x})}{(\nu(\|\mathbf{x}\|^2) - \|\nu(\mathbf{x})\|^2)^{1/2}} \right].$$

Moreover, the invariant measure $\hat{\mu}$ is determined totally by the marginal of $\tilde{\mu}$ with respect to the projection onto the x -axis.

Let us note that each point mass measure is a fixed point of T . Also, if ν is not concentrated at $E(\mathbf{x}) = \nu(\mathbf{x}) = \mathbf{m}$, then

$$(2.12) \quad \sigma^2 = E(\|\mathbf{x} - E(\mathbf{x})\|^2) = \nu(\|\mathbf{x}\|^2) - \|\nu(\mathbf{x})\|^2 > 0.$$

In this case, $\tau(A) = \nu(\sigma A + \mathbf{m})$, is a probability measure with momentum zero and energy one. If $\{T^n(\tau)\}_{n=1}^{\infty}$ converges weakly to $\hat{\mu}$, then by the commutivity properties of T , $T^n\tau((\cdot - \mathbf{m})/\sigma) = T^n(\nu)$ would converge weakly to $\hat{\mu}((\cdot - \nu(\mathbf{x})))/(\nu(\|\mathbf{x}\|^2) - \|\nu(\mathbf{x})\|^2)^{1/2}$. Therefore, to prove the theorem, we only need to prove (2.11) under the conditions that $\nu(\mathbf{x}) = 0$ and $\nu(\|\mathbf{x}\|^2) = 1$.

Our strategy is to first obtain a recursion formula for the moments of $T^n(\nu)$. Second, to show convergence to these moments and finally, to show that the limits are the moments of a unique element of $Prob(\mathbf{R}^3)$. This is the same strategy employed in [2].

3. THE MOMENT RECURSION FORMULAS.

Temporarily fix a probability measure ν on \mathbf{R}^3 with $\nu(\mathbf{x}) = 0$, $\nu(\|\mathbf{x}\|^2) = 1$. Let \mathbf{Z}_+ be the set of all non-negative integers. For each multi-index $\mathbf{k} = [k_1, k_2, k_3] \in \mathbf{Z}_+^3$, consider the mixed moment of order \mathbf{k} of the n th iterate of ν under T :

$$(3.1) \quad m_{n,\mathbf{k}} = \int_{\mathbf{R}^3} \mathbf{x}^{\mathbf{k}} dT^n \nu(\mathbf{x}) = \int_{\mathbf{R}^3} \prod_{i=1}^3 x_i^{k_i} dT^n \nu(\mathbf{x}).$$

We will first find a formula relating the moments of the $(n+1)$ th iterate to those of the n th. From (2.1), we have

$$(3.2) \quad \begin{aligned} m_{n+1,\mathbf{k}} &= \int_{\mathbf{R}^3} \mathbf{x}^{\mathbf{k}} dT(T^n(\nu)(\mathbf{x})) \\ &= \int \dots \int \left[\frac{\mathbf{S}_1}{3} + \sqrt{\lambda} \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^{1/2} \mathbf{u} \right]^{\mathbf{k}} dT^n(\nu) dT^n(\nu) d\pi_2(\mathbf{u}) d\mu(\lambda) \\ &= \int \dots \int \prod_{i=1}^3 \left[\frac{S_{1,i}}{3} + \sqrt{\lambda} \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^{1/2} u_i \right]^{k_i} dT^n(\nu) \dots d\mu(\lambda) \\ &= \int \dots \int \prod_{i=1}^3 \left[\sum_{j=0}^{k_i} \binom{k_i}{j} \sqrt{\lambda}^j \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^{i/2} u_i^j \left(\frac{S_{1,i}}{3} \right)^{k_i-j} \right] dT^n(\nu) \dots d\mu(\lambda). \end{aligned}$$

Since π_2 is the uniform distribution on the unit sphere,

$$(3.3) \quad \int_{S^2} \mathbf{u}^{\mathbf{j}} d\pi_2(\mathbf{u}) = 0,$$

unless $\mathbf{j} = [j_1, j_2, j_3]$ is even (each j_i is even). Let $[\cdot]$ denote the greatest integer function. Thus,

$$(3.4) \quad m_{n+1,\mathbf{k}} = \int \prod_{i=1}^3 \sum_{j=0}^{[k_i/2]} \binom{k_i}{2j} \lambda^j \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^j u_i^{2j} \left(\frac{S_{1,i}}{3} \right)^{k_i-2j} dT^n(\nu) \dots d\mu(\lambda).$$

Now, the weight of \mathbf{k} , $wt(\mathbf{k}) := k_1 + k_2 + k_3$. Noting that each additive term, associated with the upper summand k_i , is a polynomial in \mathbf{x}_1 , \mathbf{x}_2 , and \mathbf{x}_3 of degree k_i and with coefficients only a function of λ and \mathbf{u} , the product must be a polynomial of degree $wt(\mathbf{k})$ in the variables \mathbf{x}_1 , \mathbf{x}_2 , and \mathbf{x}_3 , with coefficients a function of λ and \mathbf{u} . Thus, integrating, we obtain an equation

$$(3.5) \quad m_{n+1, \mathbf{k}} = \sum_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3 \in \mathbf{Z}_+^3 : wt(\sum \mathbf{s}_i) = wt(\mathbf{k})} a_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3}(\mathbf{k}) m_{n, \mathbf{s}_1} m_{n, \mathbf{s}_2} m_{n, \mathbf{s}_3},$$

where the coefficients $a_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3}(\mathbf{k})$ do not depend on ν , but only on the redistribution law $\tilde{\mu}$. Also, if τ is a permutation of $\{1, 2, 3\}$, then

$$(3.6) \quad a_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3}(\mathbf{k}) = a_{\mathbf{s}_{\tau(1)}, \mathbf{s}_{\tau(2)}, \mathbf{s}_{\tau(3)}}(\mathbf{k}).$$

There is only one multi-index \mathbf{k} with $wt(\mathbf{k}) = 0$, and

$$(3.7) \quad m_{n, [0, 0, 0]} = 1, \quad n = 0, 1, 2, \dots$$

The canonical unit vectors \mathbf{e}_1 , \mathbf{e}_2 , and \mathbf{e}_3 are the only multi-indices of weight one. It is easy to check that

$$(3.8) \quad m_{n, \mathbf{e}_i} = 0, \quad n = 0, 1, 2, \dots$$

Now, according to (3.5)

$$(3.9) \quad m_{n+1, \mathbf{k}} = \sum_{I(\mathbf{k})} a_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3}(\mathbf{k}) m_{n, \mathbf{s}_1} m_{n, \mathbf{s}_2} m_{n, \mathbf{s}_3} + 3 \left(\sum_{wt(\mathbf{s}) = wt(\mathbf{k})} a_{\mathbf{s}, 0, 0}(\mathbf{k}) m_{n, \mathbf{s}} \right),$$

where the first sum is over the index set $I(\mathbf{k}) = \{(\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3) \in \mathbf{Z}_+^3 : \sum_{i=1}^3 wt(\mathbf{s}_i) = wt(\mathbf{k}), \text{ and } \forall i, wt(\mathbf{s}_i) < wt(\mathbf{k})\}$. Define

$$\gamma_n(\mathbf{k}) = \sum_{I(\mathbf{k})} a_{\mathbf{s}_1, \mathbf{s}_2, \mathbf{s}_3}(\mathbf{k}) m_{n, \mathbf{s}_1} m_{n, \mathbf{s}_2} m_{n, \mathbf{s}_3}$$

and, for $wt(\mathbf{k}) = wt(\mathbf{s})$,

$$\mathbf{A}_{\mathbf{k}, \mathbf{s}} = a_{\mathbf{s}, 0, 0}(\mathbf{k}).$$

Thus, (3.9) becomes

$$(3.10) \quad m_{n+1, \mathbf{k}} = \gamma_n(\mathbf{k}) + 3 \sum_{wt(\mathbf{s}) = wt(\mathbf{k})} \mathbf{A}_{\mathbf{k}, \mathbf{s}} m_{n, \mathbf{s}}.$$

Or, defining $y_n(p) = \{m_{n, \mathbf{s}}\}_{wt(\mathbf{s})=p}$, for $n = 0, 1, 2, \dots$; $p = 0, 1, 2, \dots$, $\gamma_n(p) = \{\gamma_n(\mathbf{k})\}_{wt(\mathbf{k})=p}$ and $\mathbf{A}(p) = \{\mathbf{A}_{\mathbf{k}, \mathbf{s}} : wt(\mathbf{k}) = wt(\mathbf{s}) = p\}$, we have

$$(3.11) \quad y_{n+1} = \gamma_n + 3\mathbf{A}y_n,$$

for $n = 0, 1, 2, \dots$. We have suppressed the argument p in (3.11) and juxtaposition signifies matrix multiplication.

Note that it follows from (3.8) that $\gamma_n(2) = 0$.

We will need to know the entries of \mathbf{A} . In order to compute these, note that for each \mathbf{k} with weight p ,

$$(3.12) \quad \sum_{wt(\mathbf{h})=p} \mathbf{A}_{\mathbf{k}, \mathbf{h}} \mathbf{x}^{\mathbf{h}} = \int \int \left[\frac{\mathbf{x}}{3} + \sqrt{\lambda} \sqrt{2/3} \|\mathbf{x}\| \mathbf{u} \right]^{\mathbf{k}} d\pi_2(\mathbf{u}) d\mu(\lambda).$$

4. CONVERGENCE OF MOMENTS OF WEIGHT 2.

From equation (3.12), we find

$$\begin{aligned}
 \phi(\mathbf{x}, \mathbf{y}) &:= \sum_{wt(\mathbf{k})=wt(\mathbf{h})=2} \binom{2}{\mathbf{k}} \mathbf{A}_{\mathbf{k}, \mathbf{h}} \mathbf{x}^{\mathbf{h}} \mathbf{y}^{\mathbf{k}} \\
 &= \iint \sum_{wt(\mathbf{k})=2} \binom{2}{\mathbf{k}} \mathbf{y}^{\mathbf{k}} \left[\frac{\mathbf{x}}{3} + \left(\frac{2\lambda}{3} \|\mathbf{x}\|^2 \right)^{1/2} \mathbf{u} \right]^{\mathbf{k}} d\pi_2(\mathbf{u}) d\mu(\lambda) \\
 (4.1) \quad &= \iint \left[\left\langle \left(\frac{\mathbf{x}}{3} + \sqrt{\frac{2\lambda}{3}} \|\mathbf{x}\| \mathbf{u} \right), \mathbf{y} \right\rangle \right]^2 d\pi_2(\mathbf{u}) d\mu(\lambda).
 \end{aligned}$$

But, since \mathbf{u} is uniformly distributed on S^2 ,

$$(4.2) \quad \phi(\mathbf{x}, \mathbf{y}) = \iint \left[\frac{\langle \mathbf{x}, \mathbf{y} \rangle}{3} + \sqrt{\lambda} \sqrt{\frac{2}{3}} \|\mathbf{x}\| \|\mathbf{y}\| u_1 \right]^2 d\pi_2(\mathbf{u}) d\mu(\lambda).$$

Since the integral of u_1 is zero,

$$(4.3) \quad \phi(\mathbf{x}, \mathbf{y}) = \frac{\langle \mathbf{x}, \mathbf{y} \rangle^2}{9} + \frac{2}{3} \|\mathbf{x}\|^2 \|\mathbf{y}\|^2 \int \lambda d\mu(\lambda) \int u_1^2 d\pi_2(\mathbf{u}).$$

Of course,

$$(4.4) \quad \int u_1^2 d\pi_2(\mathbf{u}) = \frac{1}{4\pi} \int_0^{2\pi} \int_0^\pi \cos^2 \theta \sin^3 \phi d\phi d\theta = \frac{1}{3}.$$

Thus, from (2.5) we get

$$(4.5) \quad \phi(\mathbf{x}, \mathbf{y}) = \frac{\langle \mathbf{x}, \mathbf{y} \rangle^2}{9} + \frac{2}{27} \|\mathbf{x}\|^2 \|\mathbf{y}\|^2.$$

The definition of ϕ and (4.5) yield the following matrix for $\mathbf{A} = \mathbf{A}(2)$:

$$\begin{bmatrix} \frac{5}{27} & \frac{2}{27} & \frac{2}{27} & 0 & 0 & 0 \\ \frac{2}{27} & \frac{5}{27} & \frac{2}{27} & 0 & 0 & 0 \\ \frac{2}{27} & \frac{2}{27} & \frac{5}{27} & 0 & 0 & 0 \\ 0 & 0 & 0 & \frac{3}{27} & 0 & 0 \\ 0 & 0 & 0 & 0 & \frac{3}{27} & 0 \\ 0 & 0 & 0 & 0 & 0 & \frac{3}{27} \end{bmatrix}$$

where the indices of the rows listed from top to bottom and the columns from left to right is the following sequence: $[2, 0, 0]$, $[0, 2, 0]$, $[0, 0, 2]$, $[1, 1, 0]$, $[1, 0, 1]$, $[0, 1, 1]$. Formula (3.8) implies $\gamma_n(2) = 0$, for $n = 1, 2, 3, \dots$. Thus,

$$(4.6) \quad \mathbf{y}_{n+1}(2) = (3\mathbf{A})^{n+1}(\mathbf{y}_0).$$

Inspection of the block matrix $3\mathbf{A}$ shows that the upper 3×3 block matrix has largest eigenvalue 1 and corresponding eigenvector $[1, 1, 1]$ and eigenvalue $1/3$ of multiplicity 2. The lower 3×3 diagonal matrix of $3\mathbf{A}$ has eigenvalue $1/3$ of multiplicity 3. Let $\mathbf{e} = 1/\sqrt{3}[1, 1, 1, 0, 0, 0]$. Then by this analysis

$$(4.7) \quad \lim_{n \rightarrow \infty} 3\mathbf{A}^{n+1}(\mathbf{y}_0) = \langle \mathbf{y}_0, \mathbf{e} \rangle \mathbf{e}.$$

Since $\langle \mathbf{y}_0, e \rangle = \nu(\|\mathbf{x}\|^2)/\sqrt{3}$, we have

$$(4.8) \quad \lim_{n \rightarrow \infty} m_{n, 2e_i} := m_{2e_i} = \frac{\text{Energy}}{3},$$

and, if $wt(\mathbf{k}) = 2$, and \mathbf{k} has an odd component, then $\lim_{n \rightarrow \infty} m_{n, \mathbf{k}} := m_{\mathbf{k}} = 0$.

5. CONVERGENCE OF HIGHER ORDER MIXED MOMENTS

Define a stochastic process $\{\mathbf{X}_n\}_{n=0}^{\infty}$ on \mathbf{R}^3 by letting $\{\lambda_n\}_{n=1}^{\infty}$ be a sequence of independent random variables all distributed as λ , $\{\mathbf{u}_n\}_{n=1}^{\infty}$ be a sequence of independent random vectors all distributed as \mathbf{u} so that $\{\lambda_n, \mathbf{u}_n\}_{n=1}^{\infty}$ forms an independent family and set

$$(5.1) \quad \mathbf{X}_{n+1} = \frac{\mathbf{X}_n}{3} + \sqrt{\frac{2\lambda_{n+1}}{3}} \|\mathbf{X}_n\| \mathbf{u}_{n+1}, \quad n = 1, 2, 3, \dots$$

Let $E_{\mathbf{x}_0}$ be the expectation operator where the process starts with $\mathbf{X}_0 = \mathbf{x}_0$, a.s. Also, let $W(p) = \{\mathbf{s} \in \mathbf{Z}_+^3 : wt(\mathbf{s}) = p\}$.

Lemma 5.1. *Let $\mathbf{x}_0 \in \mathbf{R}^3$, $p = wt(\mathbf{k})$, and $\mathbf{A} = A(p)$. Then*

$$(5.2) \quad E_{\mathbf{x}_0}(\mathbf{X}_n^{\mathbf{k}}) = \left(\mathbf{A}^n \mathbf{x}_0^{W(p)} \right) (\mathbf{k}), \quad n = 0, 1, 2, \dots$$

where

$$(5.3) \quad \mathbf{x}_0^{W(p)} := [\mathbf{x}_0^{\mathbf{s}}]_{\mathbf{s} \in W(p)}.$$

Clearly (5.2) is true if $n = 0$. If (5.2) holds for n , we have

$$\begin{aligned} E_{\mathbf{x}_0}(\mathbf{X}_{n+1}^{\mathbf{k}}) &= E_{\mathbf{x}_0} \left[\left(\frac{\mathbf{X}_n}{3} + \sqrt{\frac{2\lambda_{n+1}}{3}} \|\mathbf{X}_n\| \mathbf{u}_{n+1} \right) \right] \\ &= E_{\mathbf{x}_0} \left[E_{\mathbf{x}_0} \left[\left(\frac{\mathbf{X}_n}{3} + \sqrt{\frac{2\lambda_{n+1}}{3}} \|\mathbf{X}_n\| \mathbf{u}_{n+1} \right)^{\mathbf{k}} \mid \mathcal{F}_n \right] \right], \end{aligned}$$

where \mathcal{F}_n is the σ -algebra generated by $\{\lambda_j, u_j \mid j < n\}$.

According to (3.12),

$$E_{\mathbf{x}_0}(\mathbf{X}_{n+1}^{\mathbf{k}}) = E_{\mathbf{x}_0} \left[\sum_{\mathbf{s} \in W(p)} \mathbf{A}_{\mathbf{k}, \mathbf{s}} \mathbf{X}_n^{\mathbf{s}} \right],$$

and by the induction hypothesis

$$= \sum_{\mathbf{s} \in W(p)} \mathbf{A}_{\mathbf{k}, \mathbf{s}} \left[\mathbf{A}^n \mathbf{x}_0^{W(p)} \right] (\mathbf{s}).$$

Or,

$$E_{\mathbf{x}_0}(\mathbf{X}_{n+1}^{\mathbf{k}}) = \left(\mathbf{A}^{n+1} \mathbf{x}_0^{W(p)} \right) (\mathbf{k}).$$

Proof. Lemma 5.2. $\mathbf{R}^{W(p)} = \text{span}\{\mathbf{x}^{W(p)} \mid \mathbf{x} \in \mathbf{R}^3\} := \mathbf{L}$.

Let $\mathbf{v} = \{v_{\mathbf{k}}\}_{\mathbf{k} \in W(p)}$ be such that $\langle \mathbf{v}, \mathbf{x}^{W(p)} \rangle = 0$, for every $\mathbf{x} \in \mathbf{R}^3$. Then $\sum_{\mathbf{k} \in W(p)} v_{\mathbf{k}} \mathbf{x}^{\mathbf{k}}$ is a polynomial which is identically 0. Therefore, all of its coefficients are zero. So, \mathbf{v} is the zero vector. This means $L = \mathbf{R}^{W(p)}$. ■

In order to study the behavior of the iterates of $3\mathbf{A}(p)$ for $p > 2$, let us make the following notations. Define

$$\begin{aligned} H(\alpha) &:= E \left[\left\| \frac{\mathbf{e}_1}{3} + \sqrt{\frac{2}{3}} \sqrt{\lambda} \mathbf{u} \right\|^{2\alpha} \right] \\ (5.4) \quad &= E \left[\left(\frac{1}{9} + 2 \frac{\sqrt{\lambda}}{3} \sqrt{\frac{2}{3}} u_1 + \frac{2}{3} \lambda \right)^\alpha \right]. \end{aligned}$$

We have

$$(5.5) \quad H(1) = \frac{1}{3}.$$

Also, since $0 \leq \lambda \leq \frac{2}{3}$, $H(\alpha)$ is decreasing. Thus,

$$(5.6) \quad H(\alpha) < \frac{1}{3}, \text{ if } 1 < \alpha.$$

Proof. Lemma 5.3. Let C_p be the cardinality of $W(p)$. Then

$$(5.7) \quad \|(3\mathbf{A})^n \mathbf{x}_0^{W(p)}\| \leq \sqrt{C_p} \|\mathbf{x}_0\|^p (3H(p/2))^n.$$

From equation (5.2), we have

$$(5.8) \quad \left| \mathbf{A}^n \left(\mathbf{x}_0^{W(p)} \right) (\mathbf{k}) \right| \leq E_{\mathbf{x}_0} (|\mathbf{X}_n^{\mathbf{k}}|) \leq E_{\mathbf{x}_0} (\|\mathbf{X}_n\|^p).$$

Setting $\alpha = p/2$,

$$\begin{aligned} E_{\mathbf{x}_0} (\|\mathbf{X}_n\|^{2\alpha}) &= E_{\mathbf{x}_0} [\langle \mathbf{X}_n, \mathbf{X}_n \rangle^\alpha] \\ &= E_{\mathbf{x}_0} \left[\left(\frac{\|\mathbf{X}_{n-1}\|^2}{9} + \sqrt{\frac{8\lambda_n}{27}} \|\mathbf{X}_{n-1}\| \langle \mathbf{u}_n, \mathbf{X}_{n-1} \rangle + \frac{2}{3} \lambda_n \|\mathbf{X}_{n-1}\|^2 \right)^\alpha \right] \\ &= E_{\mathbf{x}_0} \|\mathbf{X}_{n-1}\|^{2\alpha} \left(\frac{1}{9} + \frac{2}{3} \sqrt{\frac{2}{3}} \sqrt{\lambda_n} \left\langle \mathbf{u}_n, \frac{\mathbf{X}_{n-1}}{\|\mathbf{X}_{n-1}\|} \right\rangle + \frac{2}{3} \lambda_n \right)^\alpha. \end{aligned}$$

Expanding the last expression in terms of conditional expectations, we have

$$E_{\mathbf{x}_0} \left[\|\mathbf{X}_{n-1}\|^{2\alpha} E_{\mathbf{x}_0} \left[\left(\frac{1}{9} + \frac{2}{3} \sqrt{\frac{2}{3}} \sqrt{\lambda_n} \left\langle \mathbf{u}_n, \frac{\mathbf{X}_{n-1}}{\|\mathbf{X}_{n-1}\|} \right\rangle + \frac{2}{3} \lambda_n \right)^\alpha \mid \mathcal{F}_{n-1} \right] \right].$$

Now, \mathbf{u}_n is uniformly distributed and independent of \mathbf{X}_{n-1} ; so we can replace $\mathbf{X}_{n-1}/\|\mathbf{X}_{n-1}\|$ by \mathbf{e}_1 . Also, the function $\left(\frac{1}{9} + \frac{2}{3} \sqrt{\frac{2}{3}} \sqrt{\lambda_n} \langle \mathbf{u}_n, \mathbf{e}_1 \rangle + \frac{2}{3} \lambda_n \right)^\alpha$ is independent of \mathcal{F}_{n-1} and has the same expected value as $\left(\frac{1}{9} + \frac{2}{3} \sqrt{\frac{2}{3}} \sqrt{\lambda} \langle \mathbf{u}, \mathbf{e}_1 \rangle + \frac{2}{3} \lambda \right)^\alpha$, which equals $\left\| \frac{\mathbf{e}_1}{3} + \sqrt{\frac{2}{3}} \sqrt{\lambda} \mathbf{u} \right\|^{2\alpha}$. Thus,

$$(5.9) \quad E_{\mathbf{x}_0} (\|\mathbf{X}_n\|^{2\alpha}) = E_{\mathbf{x}_0} [\|\mathbf{X}_{n-1}\|^{2\alpha}] H(\alpha).$$

By recursion on (5.9), we have

$$(5.10) \quad E_{\mathbf{x}_0} (\|\mathbf{X}_n\|^{2\alpha}) = \|\mathbf{x}_0\|^{2\alpha} (H(\alpha))^n.$$

Therefore, from (5.8), (5.10) and the fact that if $|z(\mathbf{k})| \leq L$, for $\mathbf{k} \in W(p)$, then $\|z\| \leq \sqrt{C_p}L$, we have

$$\|(3\mathbf{A})^n \mathbf{x}_0^{W(p)}\| \leq \sqrt{C_p} \|\mathbf{x}_0\|^p (3H(p/2))^n.$$

From Lemmas 5.2 and 5.3, we obtain

Proof. Lemma 5.4. There is a constant $D = D_p$ such that if $\mathbf{v} \in R^{W(p)}$

$$(5.11) \quad \|(3\mathbf{A})^n(\mathbf{v})\| \leq D \|\mathbf{v}\|^p (3H(p/2))^n.$$

We investigate the convergence properties of \mathbf{y}_n . Assume $p > 2$ and for each $q < p$, $\lim_{n \rightarrow \infty} \mathbf{y}_n(q) = \mathbf{y}(q)$. In the previous sections we have demonstrated the convergence of $\mathbf{y}_n(0)$, $\mathbf{y}_n(1)$, and $\mathbf{y}_n(2)$. Since $p > 2$, $3H(p/2) < 1$; and, it follows from Lemma 5.4 that the spectral radius of $3\mathbf{A} \leq 3H(p/2) < 1$. This means that the operator $(I - 3\mathbf{A})^{-1}$ exists and is equal to $\sum_{j=0}^{\infty} (3\mathbf{A})^j$. Next, a check of the definition of $\gamma_j(p)$ shows that there is a continuous function f_p such that for all j ,

$$\gamma_j(p) = f_p(\mathbf{y}_j(q); q < p).$$

So, by our assumption

$$(5.12) \quad \lim_{j \rightarrow \infty} \gamma_j(p) = f_p(\mathbf{y}(q); q < p) := \gamma(p).$$

We claim now that

$$(5.13) \quad \lim_{n \rightarrow \infty} \mathbf{y}_n(p) = (I - 3\mathbf{A})^{-1}(\gamma).$$

By recursion on (3.11), we have

$$(5.14) \quad \mathbf{y}_{n+1} = \gamma_n + 3\mathbf{A}\gamma_{n-1} + (3\mathbf{A})^2\gamma_{n-2} + \dots + (3\mathbf{A})^n\gamma_1 + (3\mathbf{A})^{n+1}\mathbf{y}_0.$$

We have

$$(5.15) \quad \|\mathbf{y}_{n+1} - (I - 3\mathbf{A})^{-1}(\gamma)\| \leq \|\gamma_n - \gamma\| + \|3\mathbf{A}(\gamma_{n-1} - \gamma)\| + \dots$$

$$(5.16) \quad + \|(3\mathbf{A})^n(\gamma_1 - \gamma)\| + \|(3\mathbf{A})^{n+1}(\mathbf{y}_0 - \gamma)\|$$

$$+ \left\| \sum_{j=n+2}^{\infty} (3\mathbf{A})^j(\gamma) \right\|.$$

For each n ,

$$(5.17) \quad \|\mathbf{y}_{n+1} - (I - 3\mathbf{A})^{-1}(\gamma)\| \leq D_j \sum_{j=1}^{\infty} c_{n,j} (3\Phi(p/2))^j,$$

where

$$c_{n,j} = \begin{cases} \|\gamma_{n-j+1} - \gamma\| & \text{if } j \leq n, \\ \|\mathbf{y}_0 - \gamma\| & \text{if } j = n+1 \\ \|\gamma\| & \text{if } j > n+1 \end{cases}.$$

Since the $c_{n,j}$'s are uniformly bounded and $\lim_{n \rightarrow \infty} c_{n,j} = 0$, we have

$$(5.18) \quad \lim_{n \rightarrow \infty} \mathbf{y}_{n+1} = (I - 3\mathbf{A})^{-1}(\gamma).$$

6. EXISTENCE AND UNIQUENESS OF A STABLE DISTRIBUTION.

In the preceding two sections we showed that there are numbers $m_{\mathbf{k}}$, $\mathbf{k} \in \mathbf{Z}_+^3$ such that if ν is an initial distribution of velocity with finite moments of all orders, then

$$(6.1) \quad \lim_{n \rightarrow \infty} m_{n, \mathbf{k}} = m_{\mathbf{k}}.$$

It is easy to see that $\{T^n(\nu)\}_{n=1}^{\infty}$ is weakly conditionally compact. Therefore, there is some probability measure $\hat{\mu}$ on \mathbf{R}^3 such that

$$(6.2) \quad m_{\mathbf{k}} = \int_{\mathbf{R}^3} \mathbf{x}^{\mathbf{k}} d\hat{\mu}(\mathbf{x}),$$

for $\mathbf{k} \in \mathbf{Z}_+^3$.

We will show that

$$(6.3) \quad \sum_{p=1}^{\infty} \left(\sum_{i=1}^3 m_{2pe_i} \right)^{-1/2p} = +\infty.$$

It follows from (6.3) that there is only one probability measure $\hat{\mu}$ with moments $m_{\mathbf{k}}$. See [4]. It also follows that $\nu, T\nu, T^2\nu, \dots$ converges weakly to $\hat{\mu}$. Finally, since T commutes with rotations, $\hat{\mu}$ is invariant under rotations. So, $\hat{\mu}$ is radially symmetric.

In order to prove (6.3), we will show that there are positive constants L and C that for all p ,

$$(6.4) \quad b_p := \hat{\mu}(\|\mathbf{x}\|^p) \leq CL^p p!,$$

for $p = 0, 1, 2, \dots$

Since $\sum_{i=1}^{\infty} m_{2pe_i} \leq b_{2p}$, for $p = 0, 1, 2, \dots$, (6.3) follows from (6.4).

In order to simplify the presentation of the argument for inequality 6.4, we will suppress the measures with respect to which the integrands are being evaluated.

We have

$$b_p = \hat{\mu}(\|\mathbf{x}\|^p) = \int \dots \int \left\| \frac{\mathbf{S}_1}{3} + \sqrt{\lambda} \sqrt{S_2 - \frac{\|\mathbf{S}_1\|^2}{3}} \mathbf{u}_0 \right\|^p.$$

Let $\tau = \text{sign}(\langle \mathbf{S}_1, \mathbf{u}_0 \rangle)$. So,

$$b_p \leq \int \dots \int \left\| \frac{\mathbf{S}_1}{3} + \tau \sqrt{\lambda} \sqrt{S_2 - \frac{\|\mathbf{S}_1\|^2}{3}} \mathbf{u}_0 \right\|^p.$$

Because τ simply flips \mathbf{u}_0 to the hemisphere with pole \mathbf{S}_1 if $\langle \mathbf{S}_1, \mathbf{u}_0 \rangle < 0$, it follows that if we replace $\sqrt{S_2 - \|\mathbf{S}_1\|^2/3}$ by a larger function, then the integral is larger. Now,

$$(6.5) \quad S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \leq \frac{2}{3} (\|\mathbf{x}_1\| + \|\mathbf{x}_2\| + \|\mathbf{x}_3\|)^2.$$

So,

$$\begin{aligned}
b_p &\leq \int \dots \int \left\| \frac{\mathbf{S}_1}{3} + \tau \sqrt{\frac{2\lambda}{3}} (\|\mathbf{x}_1\| + \|\mathbf{x}_2\| + \|\mathbf{x}_3\|) \mathbf{u}_0 \right\|^p \\
&= \int \dots \int \left\| \sum_{i=1}^3 \frac{\mathbf{x}_i}{3} + \tau \sqrt{\frac{2\lambda}{3}} \|\mathbf{x}_i\| \mathbf{u}_0 \right\|^p \\
&\leq \int \dots \int \left(\sum_{i=1}^3 \left\| \frac{\mathbf{x}_i}{3} + \tau \sqrt{\frac{2\lambda}{3}} \|\mathbf{x}_i\| \mathbf{u}_0 \right\| \right)^p.
\end{aligned}$$

Thus, setting $\mathbf{u}_i = \mathbf{x}_i / \|\mathbf{x}_i\|$, we have

$$\begin{aligned}
(6.6) \quad b_p &\leq \int \dots \int \sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} \prod_{i=1}^3 \left\| \frac{\mathbf{x}_i}{3} + \sqrt{\frac{2\lambda}{3}} \|\mathbf{x}_i\| \mathbf{u}_0 \right\|^{j_i} \\
&= \int \dots \int \sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} \prod_{i=1}^3 \left[\|\mathbf{x}_i\|^{j_i} \left\| \frac{\mathbf{u}_i}{3} + \tau \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0 \right\|^{j_i} \right].
\end{aligned}$$

The biggest $\tau \sqrt{\frac{2\lambda}{3}}$ can be is $2/3$ and \mathbf{u}_0 is independent \mathbf{u}_i . So, almost surely $\|\frac{\mathbf{u}_i}{3} + \tau \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0\| < 1$. We replace this function by 1 in (6.6) except when $j = p\mathbf{e}_i$. Thus,

$$(6.7) \quad b_p \leq \int \dots \int \left[\sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} \prod_{i=1}^3 \|\mathbf{x}_i\|^{j_i} + \sum_{i=1}^3 \left(\|\mathbf{x}_i\|^p \left\| \frac{\mathbf{u}_i}{3} + \tau_i \mathbf{u}_0 \right\|^p - \|\mathbf{x}_i\|^p \right) \right],$$

where $\tau_i = \text{sign} \langle \mathbf{u}_i, \mathbf{u}_0 \rangle$. Or,

$$(6.8) \quad b_p \leq \sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} b_{j_1} b_{j_2} b_{j_3} + \int 3 \|\mathbf{x}_i\|^p \left\| \frac{\mathbf{u}_i}{3} + \tau_i \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0 \right\|^p - 3b_p.$$

Consider the middle term of (6.8). Let $\tau^* = \text{sign} \langle \mathbf{e}_1, \mathbf{u}_0 \rangle$. We have

$$\begin{aligned}
(6.9) \quad b_p &\leq \sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} b_{j_1} b_{j_2} b_{j_3} \\
&\quad + 3b_p E \left\| \frac{\mathbf{e}_1}{3} + \tau^* \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0 \right\|^p - 3b_p.
\end{aligned}$$

Or, setting $E_p = E \|\mathbf{e}_1/3 + \tau^* \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0\|^p$,

$$(6.10) \quad (4 - 3E_p) b_p \leq \sum_{wt(\mathbf{j})=p} \binom{p}{\mathbf{j}} b_{j_1} b_{j_2} b_{j_3}.$$

Again, $\|\frac{\mathbf{e}_1}{3} + \tau^* \sqrt{\frac{2\lambda}{3}} \mathbf{u}_0\| < 1$ almost surely. Thus, $\lim_{p \rightarrow \infty} E_p = 0$. Fix p_0 such that if $p \geq p_0$, then $3E_p < 1$. Define $\{B_p\}_{p=1}^\infty$ by $B_p = b_p$, if $p < p_0$. and if $p \geq p_0$,

then by recursion,

$$(6.11) \quad B_p := \left(\frac{1}{1 - 3E_{p_0}} \right) \sum_{\substack{wt(\mathbf{j})=p \\ \mathbf{j} \neq p\mathbf{e}_i}} \binom{p}{\mathbf{j}} B_{j_1} B_{j_2} B_{j_3}.$$

Now, it follows by induction that $b_p \leq B_p$, $p = 0, 1, 2, \dots$.

Consider the formal sum

$$(6.12) \quad \begin{aligned} \Phi(t) &:= \sum_{p=0}^{\infty} \frac{B_p t^p}{p!} \\ &= \sum_{p=0}^{p_0-1} \frac{B_p t^p}{p!} + \sum_{p=p_0}^{\infty} \frac{B_p t^p}{p!}. \end{aligned}$$

We have

$$(6.13) \quad (4 - 3E_{p_0})\Phi(t) = (4 - 3E_{p_0}) \sum_{p=0}^{p_0-1} \frac{B_p t^p}{p!} + \sum_{p=p_0}^{\infty} \sum_{\substack{\mathbf{j}: \\ wt(\mathbf{j})=p}} \frac{B_{j_1} t^{j_1}}{j_1!} \frac{B_{j_2} t^{j_2}}{j_2!} \frac{B_{j_3} t^{j_3}}{j_3!}.$$

Therefore,

$$(6.14) \quad (4 - 3E_{p_0})\Phi(t) - \Phi^3(t) = p(t),$$

where $p(t)$ is a polynomial. Set

$$(6.15) \quad g(z) = (4 - 3E_{p_0})z - z^3.$$

So,

$$(6.16) \quad g(\Phi(t)) = p(t).$$

Now, $g'(1) = 1 - 3E_{p_0} > 0$. Thus, g^{-1} is analytic in a neighborhood of $g(1) = p(0)$. So, $\Phi(t) = g^{-1}(p(t))$ is analytic in a neighborhood of 0. Since the coefficients of the power series expansion of Φ about 0 are $B_p/p!$, there is a constant M such that

$$(6.17) \quad \overline{\lim}_{p \rightarrow \infty} \left[\frac{B_p}{p!} \right]^{1/p} = M < \infty.$$

Therefore, there is a constant C such that for all p ,

$$B_p \leq C(M + 1)^p p!$$

which establishes (6.4). Thus we have established Theorem 2.1 except for the last sentence. But that follows from the fact the moment recursion formulas are only functions of the moments of the marginals. See Remark 7.1 and Corollary 7.4

Interestingly enough, there is a partial converse to the Theorem 2.1.

Theorem 6.1. *If $\hat{\mu}$ is the unique fixed point of both $T_{\tilde{\mu}_1}$ and $T_{\tilde{\mu}_2}$ iff $\mu_1 = \mu_2$ (i.e., $\tilde{\mu}_1$ and $\tilde{\mu}_2$ have equal marginals.).*

That the invariant measure It is enough to show that knowing the moments of $\hat{\mu}$ and that $\hat{\mu} = T_{\tilde{\mu}}(\hat{\mu})$ uniquely determines the moments of μ , the marginal of $\tilde{\mu}$,

since μ is defined on the bounded interval $[0, \frac{2}{3}]$. To this end note that (2.1) implies $\hat{\mu}(\|\mathbf{x}\|^{2n})$ equals

$$\iiint \left[\left\| \frac{\mathbf{S}_1}{3} \right\|^2 + 2\sqrt{\lambda} \left\langle \frac{\mathbf{S}_1}{3}, \mathbf{u} \right\rangle \sqrt{S_2 - \frac{\|\mathbf{S}_1\|^2}{3}} + \frac{\mathbf{S}_1}{3} + \lambda \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right) \right]^n d\hat{\mu}^3 d\pi_2 d\mu(\lambda),$$

which equals upon applying an obvious multinomial expansion

$$\sum_{i+j+k=n} \binom{n}{i, j, k} \iiint \left\| \frac{\mathbf{S}_1}{3} \right\|^{2i} \left\langle \frac{\mathbf{S}_1}{3}, \mathbf{u} \right\rangle^j \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^{k+\frac{j}{2}} \lambda^{k+\frac{j}{2}} d\hat{\mu}^3 d\pi_2 d\mu(\lambda).$$

In the last sum, if j is odd, then the term $\left\langle \frac{\mathbf{S}_1}{3}, \mathbf{u} \right\rangle^j$ forces the term to be zero, so only whole powers of λ appear in the sum. Further, there is only one term which has λ^n , namely $i = j = 0; k = n$, and the coefficient $\int \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^n d\hat{\mu}^3$ is positive. Thus, the moment $\int \lambda^n d\mu(\lambda)$ can be written as the fraction

$$\frac{\hat{\mu}(\|\mathbf{x}\|^{2n}) - \sum_{i,j,k} \binom{n}{i,j,k} \iiint \left\| \frac{\mathbf{S}_1}{3} \right\|^{2i} \left\langle \frac{\mathbf{S}_1}{3}, \mathbf{u} \right\rangle^j \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^{k+\frac{j}{2}} \lambda^{k+\frac{j}{2}} d\hat{\mu}^3 d\pi_2 d\mu(\lambda)}{\int \left(S_2 - \frac{\|\mathbf{S}_1\|^2}{3} \right)^n d\hat{\mu}^3},$$

where the sum is over $\{(i, j, k) : k \neq n \text{ and } j \text{ even}\}$. ■

7. THE NORMAL DISTRIBUTION

In this section we show that there is a redistribution of energy law, $\tilde{\mu}$, such that the normal distribution on \mathbf{R}^3 , $N(O, I_3)$ is its attractive invariant distribution of velocity. In order to demonstrate this, we fix the following notation. Let I_n denote the $n \times n$ identity matrix, S^n the unit sphere in \mathbf{R}^{n+1} , and π_n the uniform distribution on S^n . Let A be the 3×9 matrix $[I_3, I_3, I_3]$. Let O be an isometry of \mathbf{R}^9 given by the orthogonal matrix

$$(7.1) \quad O = \begin{bmatrix} B \\ \frac{1}{\sqrt{3}} A \end{bmatrix}.$$

For example, if U, V , and W are orthogonal 3×3 matrices each with last row $(1/\sqrt{3}, 1/\sqrt{3}, 1/\sqrt{3})$, then

$$(7.2) \quad \begin{bmatrix} u_{11} & 0 & 0 & u_{12} & 0 & 0 & u_{13} & 0 & 0 \\ u_{21} & 0 & 0 & u_{22} & 0 & 0 & u_{23} & 0 & 0 \\ 0 & v_{11} & 0 & 0 & v_{12} & 0 & 0 & v_{13} & 0 \\ 0 & v_{21} & 0 & 0 & v_{22} & 0 & 0 & v_{23} & 0 \\ 0 & 0 & w_{11} & 0 & 0 & w_{21} & 0 & 0 & w_{13} \\ 0 & 0 & w_{21} & 0 & 0 & w_{22} & 0 & 0 & w_{23} \\ \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 \\ 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 \\ 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} \end{bmatrix}$$

is such a matrix O .

We will employ the three 3×9 matrices

$$(7.3) \quad \begin{aligned} M_1 &= [I_3, \mathbf{0}, \mathbf{0}] \\ M_2 &= [\mathbf{0}, I_3, \mathbf{0}] \\ M_3 &= [\mathbf{0}, \mathbf{0}, I_3]. \end{aligned}$$

Let O be an orthogonal 9×9 matrix of the form of (7.1). Let φ_0 be the map from \mathbf{R}^6 into \mathbf{R}^3 given by $\varphi_0(\mathbf{y}) = \langle \lambda_i \rangle_{i=1}^3$, where

$$(7.4) \quad \lambda_i = \left\| M_i O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} \right\|^2, \quad i = 1, 2, 3,$$

for $\mathbf{y} \in \mathbf{R}^6$, and

$$\mathbf{y} * \mathbf{0} = \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} := \begin{bmatrix} y_1 \\ y_2 \\ \vdots \\ y_6 \\ 0 \\ 0 \\ 0 \end{bmatrix}.$$

Proof. Lemma 7.1. For each O of the form of (7.1), φ_O maps S^5 onto the disk D .

Let $\mathbf{c}_1, \dots, \mathbf{c}_9$ be the column vectors of B . We have

$$(7.5) \quad O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} = \begin{bmatrix} \langle \mathbf{c}_1, \mathbf{y} \rangle \\ \vdots \\ \langle \mathbf{c}_9, \mathbf{y} \rangle \end{bmatrix}.$$

So,

$$(7.6) \quad \begin{aligned} \sum_{i=1}^3 \lambda_i &= \sum_{i=1}^3 \left\| M_i O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} \right\|^2 \\ &= \sum_{i=1}^9 \langle \mathbf{c}_i, \mathbf{y} \rangle^2. \end{aligned}$$

But, for each i , $\langle \mathbf{c}_i, \mathbf{y} \rangle = \langle \mathbf{k}_i, \mathbf{y} * \mathbf{0} \rangle$, where \mathbf{k}_i is the i^{th} column of O . Thus, (1.4)

$$\sum_{i=1}^3 \lambda_i = \|\mathbf{y} * \mathbf{0}\|^2 = 1.$$

By the orthogonality of the rows of O , we have:

$$(7.7) \quad \begin{aligned} \mathbf{c}_1 + \mathbf{c}_4 + \mathbf{c}_7 &= \mathbf{0} \\ \mathbf{c}_2 + \mathbf{c}_5 + \mathbf{c}_8 &= \mathbf{0} \\ \mathbf{c}_3 + \mathbf{c}_6 + \mathbf{c}_9 &= \mathbf{0}. \end{aligned}$$

Also, from (7.5),

$$(7.8) \quad (M_1 + M_2 + M_3) O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} = (1/\sqrt{3}) \begin{bmatrix} \langle \mathbf{c}_1 + \mathbf{c}_4 + \mathbf{c}_7, \mathbf{y} \rangle \\ \langle \mathbf{c}_2 + \mathbf{c}_5 + \mathbf{c}_8, \mathbf{y} \rangle \\ \langle \mathbf{c}_3 + \mathbf{c}_6 + \mathbf{c}_9, \mathbf{y} \rangle \end{bmatrix} = \mathbf{0}.$$

For each i , set $w_i = \frac{1}{\sqrt{\lambda_i}} M_i O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}$, if $\lambda_i \neq 0$ and $w_i = 0$ if $\lambda_i = 0$. We have (1.5)

$$\sqrt{\lambda_1} w_1 + \sqrt{\lambda_2} w_2 + \sqrt{\lambda_3} w_3 = 0.$$

Thus according to the results of section 1, φ_O maps S^5 into D . It can be checked that φ_O actually maps S^5 onto D .

For each O of the form of (7.1), let $\tilde{\mu}_0$ be the probability measure on D defined by

$$\tilde{\mu}_0(E) = \pi_5(\varphi_0^{-1}(E)).$$

Proof. Lemma 7.2. If O is of the form of (7.1), then $\tilde{\mu}_O$ is uniform measure on D .

We begin by showing that $\tilde{\mu}_{O_1} = \tilde{\mu}_{O_2}$ for any two matrices O_1, O_2 of form of (7.1). First note for any orthogonal matrix $O = \begin{bmatrix} B \\ \frac{1}{\sqrt{3}}A \end{bmatrix}$,

$$I_9 = OO^T = \begin{bmatrix} BB^T & \frac{1}{\sqrt{3}}BA^T \\ \frac{1}{\sqrt{3}}AB^T & I_3 \end{bmatrix}.$$

Thus

$$(7.9) \quad BB^T = I_6, AB^T = \mathbf{0}, \text{ and } BA^T = \mathbf{0}.$$

Let $O_i = \begin{bmatrix} B_i \\ \frac{1}{\sqrt{3}}A \end{bmatrix}$, for $i = 1, 2$. Then, $O_1 O_2^T = \begin{bmatrix} B_1 B_2^T & \frac{1}{\sqrt{3}} B_1 A^T \\ \frac{1}{\sqrt{3}} A B_2^T & I_3 \end{bmatrix} = \begin{bmatrix} B_1 B_2^T & \mathbf{0} \\ \mathbf{0} & I_3 \end{bmatrix}$ which is an orthogonal 9×9 matrix. And so $B_1 B_2^T$ is an orthogonal 6×6 matrix. Thus $O_1 O_2^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}$ has the same distribution as $\begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}$ for uniform \mathbf{y} . Easily, $\tilde{\mu}_{O_1} = \tilde{\mu}_{O_2}$.

Next we show $\tilde{\mu}_O$ is invariant under the rotation group of the disk D . Let O be chosen to be the matrix

$$(7.10) \quad \begin{bmatrix} \sqrt{\frac{2}{3}} & 0 & 0 & -\frac{1}{\sqrt{6}} & 0 & 0 & -\frac{1}{\sqrt{6}} & 0 & 0 \\ 0 & 0 & 0 & -\frac{1}{\sqrt{2}} & 0 & 0 & \frac{1}{\sqrt{2}} & 0 & 0 \\ 0 & \sqrt{\frac{2}{3}} & 0 & 0 & -\frac{1}{\sqrt{6}} & 0 & 0 & -\frac{1}{\sqrt{6}} & 0 \\ 0 & 0 & 0 & 0 & -\frac{1}{\sqrt{2}} & 0 & 0 & \frac{1}{\sqrt{2}} & 0 \\ 0 & 0 & \sqrt{\frac{2}{3}} & 0 & 0 & -\frac{1}{\sqrt{6}} & 0 & 0 & -\frac{1}{\sqrt{6}} \\ 0 & 0 & 0 & 0 & 0 & -\frac{1}{\sqrt{2}} & 0 & 0 & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 \\ 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 \\ 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} & 0 & 0 & \frac{1}{\sqrt{3}} \end{bmatrix}.$$

For $0 \leq \gamma_1, \gamma_2, \gamma_3 \leq \frac{\pi}{2}$ with $\sum_{i=1}^3 \cos^2 \gamma_i = 1$ and $0 \leq \theta_1, \theta_2, \theta_3 < 2\pi$, the vector

$$(7.11) \quad \mathbf{y} = \begin{bmatrix} \cos \gamma_1 \cos \theta_1 \\ \cos \gamma_1 \sin \theta_1 \\ \cos \gamma_2 \cos \theta_2 \\ \cos \gamma_2 \sin \theta_2 \\ \cos \gamma_3 \cos \theta_3 \\ \cos \gamma_3 \sin \theta_3 \end{bmatrix}$$

represents an arbitrary vector of S^5 . Defining $\lambda_1, \lambda_2, \lambda_3$ by (7.4), using double angle formulas

$$\lambda_1 = \frac{2}{3} \sum_{i=1}^3 \cos^2 \gamma_i \cos^2 \theta_i = \frac{1}{3} + \frac{1}{3} \sum_{i=1}^3 \cos^2 \gamma_i \cos 2\theta_i$$

and similarly

$$\begin{aligned} \gamma_2 &= \frac{1}{3} - \frac{1}{6} \sum_{i=1}^3 \cos^2 \gamma_i \cos 2\theta_i + \frac{1}{2\sqrt{3}} \sum_{i=1}^3 \cos^2 \gamma_i \sin 2\theta_i \\ \gamma_3 &= \frac{1}{3} - \frac{1}{6} \sum_{i=1}^3 \cos^2 \gamma_i \cos 2\theta_i - \frac{1}{2\sqrt{3}} \sum_{i=1}^3 \cos^2 \gamma_i \sin 2\theta_i. \end{aligned}$$

Thus the representation of $[\lambda_1, \lambda_2, \lambda_3]$ of the form of display (1.12) is

$$(7.12) \quad \begin{aligned} r \cos \theta &= \sum_{i=1}^3 \cos^2 \gamma_i \cos 2\theta_i \\ r \sin \theta &= \sum_{i=1}^3 \cos^2 \gamma_i \sin 2\theta_i. \end{aligned}$$

Multiplying (7.12) by the rotation matrix $\begin{bmatrix} \cos \psi & -\sin \psi \\ \sin \psi & \cos \psi \end{bmatrix}$ we obtain

$$(7.13) \quad \begin{aligned} r \cos(\theta - \psi) &= \sum_{i=1}^3 \cos^2 \gamma_i \cos(2\theta_i - \psi) \\ r \sin(\theta - \psi) &= \sum_{i=1}^3 \cos^2 \gamma_i \sin(2\theta_i - \psi). \end{aligned}$$

Since *given* γ_1, γ_2 , and γ_3 , the π_5 -conditional distribution of the θ_i 's are independent uniforms on $[0, 2\pi)$, it follows that (7.13) has the same distribution as (7.12). Thus, $\tilde{\mu}_O$ is rotation invariant.

Next the distribution of the variable ' $r \cos \theta$ ' of the (1.12) representation will be calculated. We choose the following parameterization of S^5 : for $\gamma \in [0, \frac{\pi}{2}]$, β , $\alpha \in [0, \pi]$, ϕ , and $\psi \in [0, 2\pi)$,

$$(7.14) \quad \mathbf{y} = \begin{bmatrix} \cos \gamma \cos \beta \\ \sin \gamma \cos \alpha \\ \cos \gamma \sin \beta \cos \phi \\ \sin \gamma \sin \alpha \cos \psi \\ \cos \gamma \sin \beta \sin \phi \\ \sin \gamma \sin \alpha \sin \psi \end{bmatrix}.$$

Using the orthogonal matrix (7.10) with the parameterization (7.14), and easy calculation using double angle formulas reveals $\lambda_1 = \frac{1}{3} + \frac{1}{3} \cos 2\gamma$. Thus $r \cos \theta$ of representation 1.11 is distributed like $\cos 2\gamma$. For parameterization (7.14), the probability density function associated with π_5 has form $C \cos^2 \gamma \sin \beta \sin^2 \gamma \sin \alpha$ for some constant C . Thus for constants C', C'' :

$$(7.15) \quad \begin{aligned} \pi_5(r \cos \theta \leq a) &= C' \int_{\frac{1}{2} \arccos(a)}^{\pi} \cos^2 \gamma \sin^2 \gamma d\gamma \\ &= \int_{-1}^a \sqrt{1-v^2} dv. \end{aligned}$$

But the only probability measure on the unit disk which is rotation invariant and which has (7.15)-distributed x -coordinate is the uniform distribution. ■

Proof. Theorem 7.3. The normal distribution, $N(\mathbf{0}, I_3)$, on \mathbf{R}^3 is the attractive invariant distribution of velocity under the redistribution of energy law which is the uniform distribution on D .

Let \tilde{v} distribution on D and v is its projection onto the x -axis. We calculate

$$(7.16) \quad \begin{aligned} \int_{\mathbf{R}^3} f dN(\mathbf{0}, I_3) &= \int_{\mathbf{R}^3} f((x_1, x_2, x_3)) dN(\mathbf{0}, I_3) \\ &= \int_0^\infty \int_{S^2} f(\sqrt{r}(w_1, w_2, w_3)) d\pi_3(\mathbf{w}) dp(r). \end{aligned}$$

Set $\mathbf{c} = \mathbf{A}w \in \mathbf{A}(S^2)$ and let $\tau(\mathbf{c})$ be the marginal distribution of \mathbf{c} . The form of O and the fact $\frac{[\mathbf{c}, \mathbf{c}, \mathbf{c}]}{3}$ is perpendicular to $O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}$ implies: for each $\mathbf{c} \in \mathbf{A}(S^2)$ we have $\mathbf{A}w = \mathbf{c}$ with $\|\mathbf{w}\| = 1$ if and only if there is some $\mathbf{y} \in S^2$ such that

$$\mathbf{w} = \frac{[\mathbf{c}, \mathbf{c}, \mathbf{c}]}{3} + \sqrt{\left(1 - \frac{\|\mathbf{c}\|^2}{3}\right)} O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}.$$

Thus,

$$[w_1, w_2, w_3] = (\mathbf{c}/3) + \sqrt{1 - \frac{\|\mathbf{c}\|^2}{3}} M_1 O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}.$$

So, (7.16) =

$$\int_0^\infty \int_{\mathbf{A}(S^2)} \int_{S^2} f \left(\sqrt{r} \left((\mathbf{c}/3) + \sqrt{\left(1 - \frac{\|\mathbf{c}\|^2}{3}\right)} M_1 O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} \right) \right) d\pi_3(\mathbf{y}) d\tau(\mathbf{c}) dp(r).$$

Now, $M_1 O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix}$ is rotation invariant on \mathbf{R}^3 . So,

$$M_1 O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} d\pi_3(\mathbf{y}) = \left\| M_1 O^T \begin{bmatrix} \mathbf{y} \\ \mathbf{0} \end{bmatrix} \right\| \mathbf{u} d\pi_3(\mathbf{y}) d\pi(\mathbf{u}).$$

Thus, (7.16) =

$$\int_0^\infty \int_{\mathbf{A}(S^2)} \int_{[0, \frac{2}{3}]} \int_{S^2} f \left[\sqrt{r} \left(\frac{\mathbf{c}}{3} + \sqrt{\lambda \left(1 - \frac{\|\mathbf{c}\|^2}{3}\right)} \mathbf{u} \right) \right] d\pi_2(\mathbf{u}) dv(\lambda) d\tau(\mathbf{c}) dp(r).$$

Putting in a dummy integral over S^5 , (7.16) =

$$\int_{[0, \frac{2}{3}]} \int_{S^2} \int_0^\infty \int_{\mathbf{A}(S^5)} \int_{S^5} f \left[\sqrt{r} \left(\frac{\mathbf{c}}{3} + \sqrt{\lambda \left(1 - \frac{\|\mathbf{c}\|^2}{3} \right)} \mathbf{u} \right) \right] d\pi_5 d\tau(\mathbf{c}) dp(r) d\pi_2(\mathbf{u}) dv(\lambda)$$

Now, $\mathbf{c} = \mathbf{A}(\mathbf{x}/\|\mathbf{x}\|) = (1/\|\mathbf{x}\|)\mathbf{A}\mathbf{x} = (1/\sqrt{S_2})\mathbf{S}_1$. So, (7.16) =

$$\int_{[0, \frac{2}{3}]} \int_{S^2} \int_{\mathbf{R}^9} f \left[\sqrt{S_2} \left(\frac{\mathbf{S}_1}{3\sqrt{S_2}} + \sqrt{\lambda \left(1 - \frac{\|\mathbf{S}_1\|^2}{3} \right)} \mathbf{u} \right) \right] dN(O, I_9) d\pi_2 dv,$$

which equals (2.1). ■

This choice \tilde{v} , uniform distribution on D , is not the only choice of energy redistribution law that attracts the Gaussian distribution of velocity. Let $Wedge = \{\mathbf{x} \in D : \exists \text{ a } \theta \in [0, \frac{\pi}{6}] \text{ and an } r \in [0, 1] \text{ s.t. equation 1.12 holds}\}$, let $\mathcal{L} = \{\alpha : Wedge \rightarrow [0, 1] : \alpha \text{ is a Borel function}\}$, and $\tilde{\mu}_\alpha$ be the measure defined through the following representation

$$\tilde{\mu}_\alpha = \int \alpha(\mathbf{x})\tau_{\mathbf{x}} + (1 - \alpha(\mathbf{x}))\tau_{Ref(\mathbf{x})} dv(x),$$

where v is uniform distribution on $Wedge$, $\tau_{\mathbf{x}}$ is the τ for \mathbf{x} in (2.8) and $Ref : D \rightarrow D$ is defined by $Ref(\mathbf{x}) = [\frac{2}{3}, \frac{2}{3}, \frac{2}{3}] - \mathbf{x}$. Let μ_α denote the projection measure of $\tilde{\mu}_\alpha$ with respect to the x -axis. The

Proof. Corollary 7.4. For all $\alpha \in \mathcal{L}$, $\mu_\alpha = v$, and therefore the attractive invariant distribution of velocity associated with $\tilde{\mu}_\alpha$ is $N(\mathbf{0}, I_3)$.

It is easy to check that $\tilde{\mu}_\alpha \circ Ref^{-1}$ is also a redistribution of energy law and that it has the same marginal projection in the x -direction as $\tilde{\mu}_\alpha$. But the measure $\frac{1}{2}(\tilde{\mu}_\alpha + \tilde{\mu}_\alpha \circ Ref^{-1}) = \tilde{v}$, uniform on D , and so the marginal of $\tilde{\mu}_\alpha$ on the x -axis is the same as v , the marginal of \tilde{v} . Apply Theorem 2.1 and Theorem 7.3. ■

Obviously $\tilde{\mu}_\alpha \neq \tilde{v}$ unless α is essentially identically $\frac{1}{2}$.

Proof. Remark 7.1. The only redistribution of energy law to produce the attractive invariant distribution of velocity law associated with the redistribution law $\delta_{[\frac{1}{3}, \frac{1}{3}, \frac{1}{3}]}$ is the energy redistribution law $\delta_{[\frac{1}{3}, \frac{1}{3}, \frac{1}{3}]}$ itself.

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